## Register Number:

DATE:

ST. JOSEPH'S COLLEGE (AUTONOMOUS), BENGALURU-27
M.Sc. PHYSICS - III SEMESTER

SEMESTER EXAMINATION: OCTOBER 2021
(Examination conducted in January-March 2022)
PH 9120 - QUANTUM MECHANICS-II
Time- 2 1/2 hrs.
Max Marks-70
This question paper has 5 printed pages and 2 parts

## Part A

## Answer any 5 questions

1. In quantum mechanics, what are the various ways Degeneracies occur? Explain
2. In the classical limit, the de-Broglie wavelength is far less than the size of the system ( $\lambda \ll x$ ). Further, in the classical limit, the total energy of the system is $E=\frac{p^{2}}{2 m}-V(x)$. With the condition that $\frac{\delta \lambda}{\lambda} \sim \frac{d \lambda}{d x}$ show that the validity of WKB approximation implies that $\delta \lambda \sim\left|\frac{m h^{2}}{p^{4}} \frac{d V}{d x}\right| \ll 1$.
3. 

(a) Write down (you don't have to derive) the perturbation equations up to the second order for a non-degenerate system described by a Hamiltonian $H^{(0)}$ and perturbed by an energy function $W=\lambda \hat{W}$ where $\lambda$ is the perturbative scale factor.
(b) Obtain the first order perturbation to the state function (use the wavefunction form, not the vectorial notation).
4. For a two fold degenerate system described by a Hamiltonian $H^{(0)}$ and perturbed by an energy function $W=\lambda \hat{W}$ where $\lambda$ is the perturbative scale factor, we know that the
equation for the first order perturbative change in energy gives us the following equation:

$$
\left|\begin{array}{cc}
\left\langle\phi_{n}^{(0) 1}\right| \hat{W}\left|\phi_{n}^{(0) 1}\right\rangle-E_{n}^{(1)} & \left\langle\phi_{n}^{(0) 1}\right| \hat{W}\left|\phi_{n}^{(0) 2}\right\rangle \\
\left\langle\phi_{n}^{(0) 2}\right| \hat{W}\left|\phi_{n}^{(0) 1}\right\rangle & \left\langle\phi_{n}^{(0) 2}\right| \hat{W}\left|\phi_{n}^{(0) 2}\right\rangle-E_{n}^{(1)}
\end{array}\right|=0 \quad \text {. What is the value of } \quad E_{n}^{(1)} ?
$$

5. Assuming that time dependent perturbation induces a transition to one of the stationary states of the unperturbed system: $|\psi(t)\rangle=\sum_{n} c_{n}(t)\left|\phi_{n}\right\rangle$, and the system to be described by a Hamiltonian $\quad H^{(0)}$ that is perturbed by an energy function $\quad W(t)=\lambda \hat{W}(t)$ with $\lambda$ being a scale factor for the perturbation, obtain an evolution equation for the coefficients after factoring out the time evolution due to the time dependent Schrodinger equation.
6. 

(a) Derive the Hamiltonian of a system of two particles in a potential dependent only on the distance between them in the center of mass frame.
(b) Express the Hamiltonian in reduced mass, and explain all terms.
7. Describe scattering by a potential with a figure. Explain the various terms and obtain the asymptotic form for the scattered wave.

## Part B

Answer any 4 questions
( $4 \times 5=20$ )
[Constants: $\mathbf{h}=6.626070 \times 10^{-34} \mathrm{~J}$ s (Planck's constant), $\mathbf{1 e V}=1.6 \times 10^{-19} \mathrm{~J}$ (electron volt to Joules), $\mathbf{c}=2.99792458 \times 10^{8} \mathrm{~m} / \mathrm{s}$ (speed of light), $1 \AA=1 \times 10^{-10} \mathrm{~m}$ (Angstrom to meters), $\mathbf{e}=1.602176 \times 10^{-19} \mathrm{C}$ (electronic charge), $\varepsilon_{0}=\quad 8.85418782 \times 10^{-12} \mathrm{~m}^{-3} \mathrm{~kg}^{-1} \mathrm{~s}^{4} \mathrm{~A}^{2} \quad$ (permittivity of free space), $\mathbf{m}_{\text {proton }}=1.672621898 \times 10^{-27} \mathrm{~kg}$ (mass of proton), $\boldsymbol{m}_{\text {electron }}=9.10938356 \times 10^{-31} \mathrm{~kg}$ (mass of electron), $\boldsymbol{m}_{\text {neutron }}=1.674927471 \times 10^{-27} \mathrm{~kg}$ (mass of neutron), $\boldsymbol{a}=5.029 \times 10^{-10} \mathrm{~m}$ (Bohr radius), $\boldsymbol{\alpha}=1 / 137$ (Fine Structure Constant), G=6.674×10-11 $\mathrm{m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2}$ (Gravitational constant), $\mathbf{M}_{\odot}=1.9891 \times 10^{30} \mathrm{~kg}$ (Solar mass), $\mathbf{R}_{\odot}=6.9 \times 10^{8} \mathrm{~m}$ (Sun's Radius), $\sigma=5.67 \times 10^{-8} \mathrm{~W} \mathrm{~m}^{-2} \mathrm{~K}^{-4}$ (Stefan-Boltzmann constant), $M_{\text {Earth }}=5.97 \times 10^{27} \mathrm{~kg}$ (Mass of Earth), $D_{\text {earth-sun }}=1.49 \times 10^{11} \mathrm{~m}$ (Earth-Sun distance), 1 inch $=2.54 \mathrm{~cm}$, 1foot=12 inches]
8. Two mutually non-interacting spin half particles (Fermions) are moving in an infinite (particle in a box) potential. The energy and wavefunction for individual particles in the infinite potential are given as: $\quad E_{n}=\frac{\hbar^{2}}{2 m} \frac{n^{2} \pi^{2}}{L^{2}}$ and $\psi_{n}=\sqrt{\frac{2}{L}} \sin \left(\frac{n \pi}{L} x\right)$ respectively. For the composite system, obtain the energy of the first excited state and write down the wavefunction (including the angular component - you can use the Table of Clebsch-Gordan coefficients available in the question paper).
9. The Hamiltonian of a two-level system described by the particle in a box potential of length
, experiences a perturbation $W=10^{-3} E_{1}^{(0)} \frac{X}{L}$ where $E_{1}^{(0)}$ is the first unperturbed eigenstate for a particle in a box. Obtain the first order change in the energies of the two levels. The energies and state functions for the unperturbed particle in a box are given as: $E_{n}^{(0)}=\frac{\hbar^{2}}{2 m} \frac{n^{2} \pi^{2}}{L^{2}} \quad$ and $\quad \psi_{n}^{(0)}=\sqrt{\frac{2}{L}} \sin \left(\frac{n \pi x}{L}\right)$ respectively.
10. What is probability of transition from the ground state $n=1$ to the first excited state $n=2$ for a perturbation of $W=10^{-3} E_{1}^{(0)} \sin \omega t$ to a system described by the particle in a box potential?
11. Use the variational method to find the optimal values for $b$ and the corresponding approximate ground state and ground state energy for a simple harmonic oscillator $V=\frac{1}{2} m \omega^{2} x^{2}$ where $m$ is the mass of the particle and $\omega$ the frequency of the oscillator. Use the trial wavefunction: $\quad \psi_{\text {trial }}(x)=\frac{A}{x^{2}+b}$. The average kinetic energy for the system works out to be: $\langle T\rangle=\frac{\hbar^{2}}{4 b m}$. You may need to use some of the standard integrals given later in the question paper.
12.
(a) Show that the ground state for a system is always a lower bound to the expectation value of energy for a system.
(b) What are the conditions on the wavefunction of the system for this to happen?
13. Using the Born approximation for scattering amplitude: $f_{k}^{(B)}(\theta, \phi)=\frac{-\mu}{2 \pi \hbar^{2}} \int V(\vec{r}) d^{3} r$ compute the scattering amplitude and differential scattering cross section of scattering by a soft-sphere potential: $V(\vec{r})=\left\{\begin{array}{ll}V_{0} & \text { if } r \leq a \\ 0 & \text { if } r>a\end{array}\right.$.

## Table of ${ }_{\text {(some) }}$ Integrals

| Gamma Function: |
| :--- |
| $\Gamma(x)=\int_{0}^{\infty} t^{x-1} e^{-t} d t$ |
| $\Gamma(n)=(n-1)!$ |
| $\Gamma\left(\frac{1}{2}+n\right)=\frac{(2 n)!}{4^{n} n!} \sqrt{ } \pi$ |

(a) $\int_{0}^{\infty} e^{-2 b t} d t=\frac{1}{2 b}$
(I) $\int \frac{t^{2}}{\left(t^{2}+b^{2}\right)^{2}} d t=\left(-\frac{t}{\left(2 b^{2}+2 t^{2}\right)}+\frac{1}{2 b} \tan ^{-1}\left(\frac{t}{b}\right)\right)$
(b) $\int_{0}^{\infty} t e^{-2 b t} d t=\frac{1}{4 b^{2}}$
(m) $\int \frac{1}{\left(t^{2}+b^{2}\right)^{3}} d t=\frac{3}{8 b^{5}}\left(\frac{5 / 3 b^{3} t+b t^{3}}{\left(b^{2}+t^{2}\right)^{2}}+\tan ^{-1}\left(\frac{t}{b}\right)\right)$
(c) $\int_{0}^{\infty} t^{2} e^{-2 b t} d t=\frac{1}{4 b^{3}}$
(n) $\int \frac{t^{2}}{\left(t^{2}+b^{2}\right)^{4}} d t=\frac{1}{16 b^{5}}\left(\frac{b t^{5}+8 / 3 b^{3} t^{3}-b^{5} t}{\left(b^{2}+t^{2}\right)^{3}}+\tan ^{-1}\left(\frac{t}{b}\right)\right)$
(d) $\int_{0}^{\infty} t^{3} e^{-2 b t} d t=\frac{3}{8 b^{4}}$
(o) $\int \frac{t^{4}}{\left(t^{2}+b^{2}\right)^{4}} d t=\frac{1}{16 b^{3}}\left(\frac{b t^{5}+8 / 3 b^{3} t^{3}-b^{5} t}{\left(b^{2}+t^{2}\right)^{3}}+\tan ^{-1}\left(\frac{t}{b}\right)\right)$
(e) $\int_{0}^{\infty} t^{4} e^{-2 b t} d t=\frac{3}{4 b^{5}}$
(p) $\int \frac{t^{6}}{\left(t^{2}+b^{2}\right)^{4}} d t=\frac{1}{16 b}\left(\frac{11 b t^{5}+40 / 3 b^{3} t^{3}-5 b^{5} t}{\left(b^{2}+t^{2}\right)^{3}}+5 \tan ^{-1}\left(\frac{t}{b}\right)\right)$
(f) $\int_{0}^{\infty} t^{5} e^{-2 b t} d t=\frac{15}{8 b^{6}}$
(q) $\int \sqrt{a / x-1} d x=x \sqrt{a / x-1}+a \tan ^{-1}(\sqrt{a / x-1})$
(g) $\int_{0}^{\infty} t^{6} e^{-2 b t} d t=\frac{45}{8 b^{7}}$
(r) $\int \sqrt{1-a x} d x=-\frac{2(1-a x)^{3 / 2}}{3 a}$
(h) $\int \frac{1}{t^{2}+b^{2}} d t=\frac{1}{b} \tan ^{-1}\left(\frac{t}{b}\right)$
(s) $\int \sqrt{1-a x^{2}} d x=\frac{1}{2} x \sqrt{1-a x^{2}}+\frac{\sin ^{-1} \sqrt{a} x}{2 \sqrt{a}}$
(i) $\int \frac{1}{\left(t^{2}+b^{2}\right)^{2}} d t=\frac{1}{2 b^{3}}\left(\frac{b t}{\left(b^{2}+t^{2}\right)}+\tan ^{-1}\left(\frac{t}{b}\right)\right)$
(t) $\int_{-\infty}^{\infty} e^{-\alpha t^{2}+i \omega t} d t=\sqrt{\frac{\pi}{\alpha}} e^{-\frac{\left(\omega^{2}\right.}{4 \alpha}}$
(j) $\int_{0}^{\infty} t^{4} e^{-\alpha^{2} t^{2}} d t=\frac{3 \sqrt{\pi}}{8 \alpha^{5}}$
(u) $\int_{0}^{\infty} t^{n} e^{-s t} d t=\frac{n!}{s^{n+1}} \quad$ (Laplace Transform)
(k) $\int \frac{1}{\left(t^{2}+b^{2}\right)^{4}} d t=\frac{1}{16 b^{7}}\left(\frac{15 t^{5} b+40 b^{3} 5^{3}+33 b^{5} t}{\left(3 t^{6}+9 b t^{4}+9 b^{3} t^{2}+b^{5}\right)}+5 \tan ^{-1}\left(\frac{t}{b}\right)\right)$

## 34. CLEBSCH-GORDAN COEFFICIENTS, SPHERICAL HARMONICS, AND $d$ FUNCTIONS



Figure 34.1: The sign convention is that of Wigner (Group Theory, Academic Press, New York, 1959), also used by Condon and Shortley (The Theory of Atomic Spectra, Cambridge Univ. Press, New York, 1953), Rose (Elementary Theory of Angular Momentum, Wiley, New York, 1957), and Cohen (Tables of the Clebsch-Gordan Coefficients, North American Rockwell Science Center, Thousand Oaks, Calif., 1974). The coefficients here have been calculated using computer programs written independently by Cohen and at LBNL.

